

Remarks on Perturbative Categorical Quantum Gravity

Jerzy Król

University of Silesia, Institute of Physics
Katowice, Poland

in cooperation with *Michael Heller*

CATEGORIES'17
Warszawa, 16-17. 11. 2017

Motivations

Feynman diagrams are typical tools of QFT formulated on flat Minkowski spacetime M^4 . They describe interactions of all quantum fields in the Standard Model, i.e. except *gravity*, in the perturbative regime.
Classically: inclusion of gravity \implies curvature on M^4 .

Motivations

Feynman diagrams are typical tools of QFT formulated on flat Minkowski spacetime M^4 . They describe interactions of all quantum fields in the Standard Model, i.e. except *gravity*, in the perturbative regime.

Classically: inclusion of gravity \implies curvature on M^4 .

However, Feynman (1950-64) tried to build the theory of quantum gravity based on the Feynman diagrams on flat M^4 .

Classical GR of Einstein: any particle, and form of energy, interact gravitationally. The Feynman's theory describes interacting (gravitationally) quantum particles via the exchange of the virtual spin-2 graviton.

Motivations

Feynman diagrams are typical tools of QFT formulated on flat Minkowski spacetime M^4 . They describe interactions of all quantum fields in the Standard Model, i.e. except *gravity*, in the perturbative regime.
Classically: inclusion of gravity \implies curvature on M^4 .

However, Feynman (1950-64) tried to build the theory of quantum gravity based on the Feynman diagrams on flat M^4 .

Classical GR of Einstein: any particle, and form of energy, interact gravitationally. The Feynman's theory describes interacting (gravitationally) quantum particles via the exchange of the virtual spin-2 graviton.

Linearized perturbations of metric undergo quantization as other fields. Whole approach is on *flat* M^4 , so it is called *covariant* QG (i.e. Lorentz invariant) and *perturbative* – there are Feynman diagrams involved.

Main features of the *perturbative covariant QG*

- i. Feynman diagrams are prime ingredients (not any classical gravity);
- ii. It is a perturbative theory;
- iii. The background is entirely flat Minkowski spacetime (hence no classical gravity = curvature);
- iv. Classical GR is restored from the tree level of the diagrams [Feynman, 1962].

General setup

- On Minkowski ST, M^4 , we have quantum fields representing electrons, photons, protons, gluons etc. They undergo quantization.
- Gravitational field is also defined on M^4 ($g_{\mu\nu}$) but it does not curve the spacetime. The linearization of $g_{\mu\nu}$, i.e. $h_{\mu\nu} = g_{\mu\nu} - \eta_{\mu\nu}$, is quantized as other fields and gives rise to gravitons $G_{\mu\nu}$.
- Gravitational interactions are described by Feynman diagrams of QFT on flat Minkowski M^4 via interchanging the gravitons.
- The Feynman diagrams are built according to the rules:
 1. vertices
 2. propagators

Problems of pQG

The Feynman's project being interesting by itself is, however, completely wrong as a physical theory. Such perturbative QG is non-renormalizable, i.e. the predictions of the theory cannot be made finite even after infinitely correcting the divergent quantities.

Problems of pQG

The Feynman's project being interesting by itself is, however, completely wrong as a physical theory. Such perturbative QG is non-renormalizable, i.e. the predictions of the theory cannot be made finite even after infinitely correcting the divergent quantities.

Currently we do not have any good (finite, renormalizable, perturbative or non-perturbative) theory of QG in dimension 4. However, pure QG is renormalizable on-shell in 1-loop, and QG with sources is non-renormalized in 1-loop [Veltman, t'Hooft, 1970's].

Problems of pQG

The Feynman's project being interesting by itself is, however, completely wrong as a physical theory. Such perturbative QG is non-renormalizable, i.e. the predictions of the theory cannot be made finite even after infinitely correcting the divergent quantities.

Currently we do not have any good (finite, renormalizable, perturbative or non-perturbative) theory of QG in dimension 4. However, pure QG is renormalizable on-shell in 1-loop, and QG with sources is non-renormalized in 1-loop [Veltman, t'Hooft, 1970's].

Perturbative Covariant QG on 4-D Minkowski spacetime, *fails*.

Problems of pQG

The Feynman's project being interesting by itself is, however, completely wrong as a physical theory. Such perturbative QG is non-renormalizable, i.e. the predictions of the theory cannot be made finite even after infinitely correcting the divergent quantities.

Currently we do not have any good (finite, renormalizable, perturbative or non-perturbative) theory of QG in dimension 4. However, pure QG is renormalizable on-shell in 1-loop, and QG with sources is non-renormalized in 1-loop [Veltman, t'Hooft, 1970's].

Perturbative Covariant QG on 4-D Minkowski spacetime, *fails*.

Instead, let us take a reasonable categorical extension of 4-D spacetime and try to carry out the Feynman project on it.

A Categorical perturbative QG

We work in a model of Synthetic Differential Geometry (SDG), i.e. some smooth topos \mathcal{T} in which the category of smooth manifolds and maps, \mathbb{M} , is represented fully and faithfully

$$s : \mathbb{M} \hookrightarrow \mathcal{T}.$$

$s(\mathbb{R}) = R$ contains various objects of infinitesimal 'numbers' ($\neq 0$)

$$D_k \rightarrow_{\mathcal{T}} R, d \in D_k | d^{k+1} = 0.$$

A Categorical perturbative QG

We work in a model of Synthetic Differential Geometry (SDG), i.e. some smooth topos \mathcal{T} in which the category of smooth manifolds and maps, \mathbb{M} , is represented fully and faithfully

$$s : \mathbb{M} \hookrightarrow \mathcal{T}.$$

$s(\mathbb{R}) = R$ contains various objects of infinitesimal 'numbers' ($\neq 0$)

$$D_k \rightarrow_{\mathcal{T}} R, d \in D_k | d^{k+1} = 0.$$

A categorical spacetime is a spacetime M^4 which 'locally' (for sufficiently fine local cover $\{U_\alpha \simeq \mathbb{R}^4\}_{\alpha \in I}$) allows the internal to \mathcal{T} reasoning, i.e.

$$\forall U_\alpha \exists U'_\alpha \subset U_\alpha (s(U'_\alpha) \simeq_{\mathcal{T}} R^4).$$

The point is that we start with M^4 in SET at macro-scale, and continue in \mathcal{T} at micro-scale. PHYSICS: in sufficiently high energies (small distances) the reasoning becomes intuitionistic and fields and spacetime live in \mathcal{T} .

Monads

In \mathcal{T} there are monads $\mathcal{M}(x), x \in s(M^4)$ in particular around $s(x)$ in \mathcal{T} such that $x \in M^4$ in SET, i.e.

$$\forall x \in M^4 (\mathcal{M}_k(x) = \{y \in s(M^4) | y - x \in D_k^4\})$$

where we refer to the local (micro)linear structure for $-$.
 $\mathcal{M}_k(x)$ is parametrized by D_k for some $k \in \mathbb{N}$. Note that

$$d^{k+1} = 0, d \in D_k.$$

Certainly in SET there are no $d \in \mathbb{R}$ that $d^{k+1} = 0, d \neq 0$.
So, we need intuitionism and topoi.

Gravity on monads

MAIN QUESTION: given monads $\mathcal{M}(x) \hookrightarrow s(M^4)$ in \mathcal{T} does gravitational field $h_{\mu\nu}(x), x \in M^4$ in SET know about them?
Can $h_{\mu\nu}$ propagate on monads?

The ANSWER: the only field propagating on monads is $h_{\mu\nu}$. It is the only field 'sensitive' on the intuitionistic structure of the spacetime at microscale.

The ANSWER: the only field propagating on monads is $h_{\mu\nu}$. It is the only field 'sensitive' on the intuitionistic structure of the spacetime at microscale.

- *Quantum* gravitational interactions take place on monads in \mathcal{T} so that the vertices with virtual gravitons live on monads.
- *Classical* gravity is described entirely on M^4 in SET.

/This discrepancy \implies problems with quantizing gravity?/

How to describe $h_{\mu\nu}$ on monads?

How to describe $h_{\mu\nu}$ on monads?

SDG: D_k is described by the spectrum, $Spec_R(W)$, of a Weil algebra W and W can be represented as $W = \mathbb{R} \oplus \mathbb{R}[\epsilon] \oplus \mathbb{R}[\epsilon^2] \oplus \dots$ (k -many ϵ terms) and the real spectrum is $\mathbb{R}[\epsilon] \oplus \mathbb{R}[\epsilon^2] \oplus \dots$. This is how D_k looks like in SET.

How to describe $h_{\mu\nu}$ on monads?

SDG: D_k is described by the spectrum, $Spec_R(W)$, of a Weil algebra W and W can be represented as $W = \mathbb{R} \oplus \mathbb{R}[\epsilon] \oplus \mathbb{R}[\epsilon^2] \oplus \dots$ (k -many ϵ terms) and the real spectrum is $\mathbb{R}[\epsilon] \oplus \mathbb{R}[\epsilon^2] \oplus \dots$. This is how D_k looks like in SET.

Certainly, D_k is not any subset of \mathbb{R} in SET (but it is a subobject of R in \mathcal{T})
 ϵ is the so called dual number, $\epsilon^{k+1} = 0$.

How to describe $h_{\mu\nu}$ on monads?

SDG: D_k is described by the spectrum, $Spec_R(W)$, of a Weil algebra W and W can be represented as $W = \mathbb{R} \oplus \mathbb{R}[\epsilon] \oplus \mathbb{R}[\epsilon^2] \oplus \dots$ (k -many ϵ terms) and the real spectrum is $\mathbb{R}[\epsilon] \oplus \mathbb{R}[\epsilon^2] \oplus \dots$. This is how D_k looks like in SET.

Certainly, D_k is not any subset of \mathbb{R} in SET (but it is a subobject of R in \mathcal{T}) ϵ is the so called dual number, $\epsilon^{k+1} = 0$.

If (quantum) gravity interacts at monads we can describe it in SET via W picture. In infinitesimal terms on D_k the gravitational coupling constant $f = \sqrt{8\pi G}$ is represented by $\sqrt{8\pi G} \cdot \epsilon$. Physics distinguishes D_2 , so ϵ is from D_2 , thus

$$(\sqrt{G} \cdot \epsilon)^3 = 0$$

and

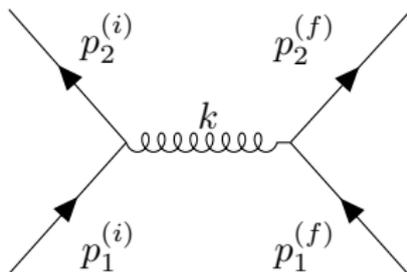
$$f_{D_2} = \sqrt{8\pi G} \cdot \epsilon$$

Since only gravity explores monads this is the only coupling with ϵ unit.

We will see that Feynman covariant QG on the categorical spacetime has very desirable features.

perturbative QG on monads and classical Newton law

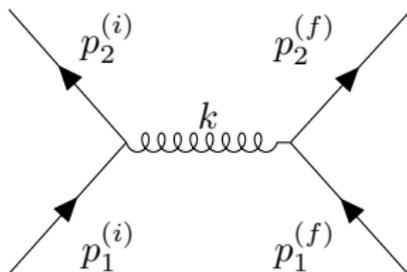
The covariant term for the interchanging of a single graviton (in the lowest tree-level) on a monad is calculated from the diagram



$$-if^2 T_{\mu\nu}^1 \frac{\mathcal{P}^{\mu\nu,\alpha\beta}}{k^2 + i} T_{\alpha\beta}^2 \delta(P_{fi}) \cdot \epsilon^2.$$

perturbative QG on monads and classical Newton law

The covariant term for the interchanging of a single graviton (in the lowest tree-level) on a monad is calculated from the diagram



$$-i f^2 T_{\mu\nu}^1 \frac{\mathcal{P}^{\mu\nu,\alpha\beta}}{k^2 + i} T_{\alpha\beta}^2 \delta(P_{fi}) \cdot \epsilon^2.$$

The Fourier transform of it (T) (in the Mandelstam coordinates, non-relativistic) reads

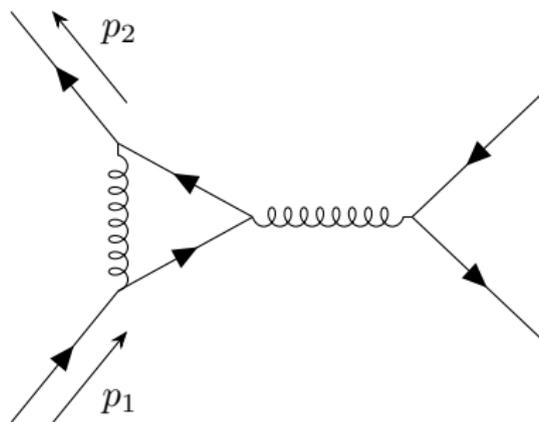
$$-\frac{1}{m_1 m_2} \int e^{-i\mathbf{k}\cdot\mathbf{r}}(T) d^3 k \cdot \epsilon^2 = \left(-\frac{f^2 m_1 m_2}{8\pi r} + \frac{1}{2} f^2 \delta^3(\mathbf{r}) \right) \cdot \epsilon^2.$$

perturbative QG on monads

Introducing other graviton vertices gives additional f_{D_2} factors and
 $\sim f_{D_2}^2 \cdot f_{D_2} = f_{D_2}^3 = 0$ – the expression vanishes.

perturbative QG on monads

Introducing other graviton vertices gives additional f_{D_2} factors and $\sim f_{D_2}^2 \cdot f_{D_2} = f_{D_2}^3 = 0$ – the expression vanishes. Below there is the 1-loop correction to the graviton – fermion vertex



where $\sim f_{D_2}^4 = 0$.

So, maximally 1-loop gravitational diagrams (these which have 2 gravity vertices) survive.

So, maximally 1-loop gravitational diagrams (these which have 2 gravity vertices) survive.

In SET: 1-loop pure gravity is renormalizable on M^4 . Full pert. QG on the ordinary spacetime is non-renormalizable.

So, maximally 1-loop gravitational diagrams (these which have 2 gravity vertices) survive.

In SET: 1-loop pure gravity is renormalizable on M^4 . Full pert. QG on the ordinary spacetime is non-renormalizable.

On the synthetic spacetime: it is expected that the pert. pure QG will be renormalizable, since it is truncated to the subset of 1-loops of QG in SET.

So, maximally 1-loop gravitational diagrams (these which have 2 gravity vertices) survive.

In SET: 1-loop pure gravity is renormalizable on M^4 . Full pert. QG on the ordinary spacetime is non-renormalizable.

On the synthetic spacetime: it is expected that the pert. pure QG will be renormalizable, since it is truncated to the subset of 1-loops of QG in SET.

Moreover, introducing gravitational vertices into known Feynman diagrams of other interactions can act as natural renormalizator. 3-gravitational vertices cancel the connected diagram of any kind.

So, maximally 1-loop gravitational diagrams (these which have 2 gravity vertices) survive.

In SET: 1-loop pure gravity is renormalizable on M^4 . Full pert. QG on the ordinary spacetime is non-renormalizable.

On the synthetic spacetime: it is expected that the pert. pure QG will be renormalizable, since it is truncated to the subset of 1-loops of QG in SET.

Moreover, introducing gravitational vertices into known Feynman diagrams of other interactions can act as natural renormalizator. 3-gravitational vertices cancel the connected diagram of any kind.

Where is classical curvature-gravity on flat Minkowski M^4 ? Do we rule out it completely?

QUESTION: can the change of SET to \mathcal{T} in the definition of a categorical spacetime, be made smooth?

QUESTION: can the change of SET to \mathcal{T} in the definition of a categorical spacetime, be made smooth?

ANSWER: Yes. However the smoothness on manifolds like $M^3 \times \mathbb{R}$ has to be non-standard (exotic) in SET, if $M^3 \times \mathbb{R}$ is to be categorical spacetime (i.e. allows for the change $\text{SET} \rightarrow \mathcal{T}$ at microscale) [JK,MH,2017].

This is the consequence of incompatibility of the real line \mathbb{R} in SET with the intuitionistic real line $R_{\mathcal{T}}$ in \mathcal{T} .

The agreement of the smoothness of spacetimes with the shift $\text{SET} \rightarrow \mathcal{T}$ is possible only in dimension 4.

The agreement of the smoothness of spacetimes with the shift $\text{SET} \rightarrow \mathcal{T}$ is possible only in dimension 4. Even though there exist exotic smooth manifolds in higher dimensions, like $S^7 \times \mathbb{R}$ or $S^{11} \times \mathbb{R}$ etc, they can not be compatible with $\text{SET} \rightarrow \mathcal{T}$.

The agreement of the smoothness of spacetimes with the shift $\text{SET} \rightarrow \mathcal{T}$ is possible only in dimension 4. Even though there exist exotic smooth manifolds in higher dimensions, like $S^7 \times \mathbb{R}$ or $S^{11} \times \mathbb{R}$ etc, they can not be compatible with $\text{SET} \rightarrow \mathcal{T}$.

In particular case \mathbb{R}^4 it is an exotic R^4 s.t. the product $\mathbb{R}^3 \times \mathbb{R}$ is not a smooth product.

The agreement of the smoothness of spacetimes with the shift $\text{SET} \rightarrow \mathcal{T}$ is possible only in dimension 4. Even though there exist exotic smooth manifolds in higher dimensions, like $S^7 \times \mathbb{R}$ or $S^{11} \times \mathbb{R}$ etc, they can not be compatible with $\text{SET} \rightarrow \mathcal{T}$.

In particular case \mathbb{R}^4 it is an exotic R^4 s.t. the product $\mathbb{R}^3 \times \mathbb{R}$ is not a smooth product. This R^4 has to be Riemann curved, and this curvature defines differential structure underlying the Minkowski spacetime.

The agreement of the smoothness of spacetimes with the shift $\text{SET} \rightarrow \mathcal{T}$ is possible only in dimension 4. Even though there exist exotic smooth manifolds in higher dimensions, like $S^7 \times \mathbb{R}$ or $S^{11} \times \mathbb{R}$ etc, they can not be compatible with $\text{SET} \rightarrow \mathcal{T}$.

In particular case \mathbb{R}^4 it is an exotic R^4 s.t. the product $\mathbb{R}^3 \times \mathbb{R}$ is not a smooth product. This R^4 has to be Riemann curved, and this curvature defines differential structure underlying the Minkowski spacetime.

The meaning of this 'exotic' curvature is physical and is under intense study currently [T.Asselmeyer-Maluga,JK,2017].

Summary

Category theory opens fundamentally new perspectives on Covariant Perturbative QG and helps understanding the task of quantizing gravity and spacetime structure at microscale.

Summary

Category theory opens fundamentally new perspectives on Covariant Perturbative QG and helps understanding the task of quantizing gravity and spacetime structure at microscale.

(-) Perturbative QG on the synthetic spacetime can be renormalizable (proof needed).

Summary

Category theory opens fundamentally new perspectives on Covariant Perturbative QG and helps understanding the task of quantizing gravity and spacetime structure at microscale.

(-) Perturbative QG on the synthetic spacetime can be renormalizable (proof needed).

(-) The unifying theory of categorical QG with other interactions can also be renormalizable in the regime of Planck energies (proof needed)

Summary

Category theory opens fundamentally new perspectives on Covariant Perturbative QG and helps understanding the task of quantizing gravity and spacetime structure at microscale.

(–) Perturbative QG on the synthetic spacetime can be renormalizable (proof needed).

(–) The unifying theory of categorical QG with other interactions can also be renormalizable in the regime of Planck energies (proof needed)

(–) The change of the category from SET to a Topos \mathcal{T} (representing SDG and \mathbb{M}) can be supported by a smooth manifold. Manifolds allowing for this exist only in dimension 4. They are exotic manifolds, like R^4 .

Summary

Category theory opens fundamentally new perspectives on Covariant Perturbative QG and helps understanding the task of quantizing gravity and spacetime structure at microscale.

(-) Perturbative QG on the synthetic spacetime can be renormalizable (proof needed).

(-) The unifying theory of categorical QG with other interactions can also be renormalizable in the regime of Planck energies (proof needed)

(-) The change of the category from SET to a Topos \mathcal{T} (representing SDG and \mathbb{M}) can be supported by a smooth manifold. Manifolds allowing for this exist only in dimension 4. They are exotic manifolds, like R^4 .

(-) Exotic R^4 can not be Riemann flat. Its curvature is physical: the curvature of $R^4 \sim$ the measured value of the CC density energy of the expanding universe.

(To explain this value is one of the biggest, still open, mysteries of physics).

a bit of formal play

Superspace Grassmann's coordinates Θ fulfill

$$\{\Theta, \Theta\} = [\Theta, \Theta] = 0 \implies \Theta^2 = 0. \quad (1)$$

Maybe we should work in yet another coordinates, $d^3 = 0$, in order to quantize gravity? (ϵ -valued physical units)

a bit of formal play

Superspace Grassmann's coordinates Θ fulfill

$$\{\Theta, \Theta\} = [\Theta, \Theta] = 0 \implies \Theta^2 = 0. \quad (1)$$

Maybe we should work in yet another coordinates, $d^3 = 0$, in order to quantize gravity? (ϵ -valued physical units)

What is the relation of $d = f \cdot \epsilon$ to Θ and $[\ , \]$ and $\{ \ , \ }$? How to express formally $d^3 = 0, d \neq 0$ by $[\ , \]$ and $\{ \ , \ }$?

a bit of formal play

Superspace Grassmann's coordinates Θ fulfill

$$\{\Theta, \Theta\} = [\Theta, \Theta] = 0 \implies \Theta^2 = 0. \quad (1)$$

Maybe we should work in yet another coordinates, $d^3 = 0$, in order to quantize gravity? (ϵ -valued physical units)

What is the relation of $d = f \cdot \epsilon$ to Θ and $[\ , \]$ and $\{ \ , \ }$? How to express formally $d^3 = 0, d \neq 0$ by $[\ , \]$ and $\{ \ , \ }$?

One way is easy, if $\Theta = d^2$ then (1) holds, and formally

$$d = \sqrt{\Theta} \dots$$

Thank You!